

Statistical Mechanics and Material Properties

By

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Statistical Mechanics and Thermodynamics

- Thermodynamics
(0th low : A.eq.B and B.eq.C \rightarrow A.eq.C)
1st low : Energy conservation
2nd low : Free energy decreases
3rd low : Entropy \rightarrow 0 at 0K

Interface essential to utilize thermodynamics

- Statistical Mechanics
Ergodic theorem : time average = space average, (in equilibrium)



Grave-stone of Boltzmann...

You may have learnt Thermodynamics

- Thermodynamics

(0th law : A.eq.B and B.eq.C \rightarrow A.eq.C)

1st law : Energy conservation

2nd law : Free energy decreases

3rd law : Entropy \rightarrow 0 at 0K

Everything is
derived from these
principles...
But !

- Definitions

- Temperature
- Free energy
- Entropy
- Equilibrium

Please remember when you
have heard them first.
How they were defined ?

**Could you understand them
as you could do in kinetics ?**

Definitions

□ Temperature

°C : water

°F : salt water



Empirical temp.

Temp. in science (physics or kinetics) is same as above ?

□ Free energy

Helmholtz

$$F = U - TS$$

Gibbs

$$G = U + PV - TS$$

Free energy is same as the energy in kinetics ?

Why they need to be defined as above ?

□ Entropy

What is this ?

□ Equilibrium

What is the definition ?

You may have learnt Thermodynamics

■ Thermodynamics

(0th law : A.eq.B and B.eq.C \rightarrow A.eq.C)

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Everything is
derived from these
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But !

■ Definitions

- Temperature
- Free energy
- Entropy
- Equilibrium

Understanding of them
(= **statistical mechanics**)
lead you to utilize
thermodynamics



Material properties

Let's define them step by step !

Temperature (first trial)

Example : Ideal gas (1)

■ Assumption

Number of molecule : N in volume : V

Velocity of molecule : $\mathbf{c} = (u, v, w) \longrightarrow c^2 = u^2 + v^2 + w^2$

probability : $F(\mathbf{c}) = f(u)f(v)f(w)$

■ Momentum change = Pressure on area A

$$\int_0^{+\infty} \left(2mu \times f(u) \frac{N}{V} A u dt \right) du = P A dt$$

$$PV = \int_0^{+\infty} (2mu \times f(u) Nu) du = mN \int_{-\infty}^{+\infty} u^2 f(u) du = Nm \langle u^2 \rangle$$

$$\langle u^2 \rangle = \langle v^2 \rangle = \langle w^2 \rangle$$

$$\langle u^2 \rangle = \frac{1}{3} \langle c^2 \rangle$$

■ Comparison with eq. of state :

$$PV = NkT$$

$$\left\langle \frac{1}{2} mc^2 \right\rangle = \frac{3}{2} kT$$

$$PV = Nm \frac{1}{3} \langle c^2 \rangle = N \frac{2}{3} \left\langle \frac{1}{2} mc^2 \right\rangle$$

■ Law of equi-partition of energy

$$\left\langle \frac{1}{2} mu^2 \right\rangle = \left\langle \frac{1}{2} mv^2 \right\rangle = \left\langle \frac{1}{2} mw^2 \right\rangle = \frac{1}{2} kT$$

Example : Ideal gas (2)

$$c^2 = u^2 + v^2 + w^2$$

■ Probability function

$$F(\mathbf{c}) = f(u)f(v)f(w)$$

taking log :

$$\ln F(\mathbf{c}) = \ln f(u) + \ln f(v) + \ln f(w)$$

diff. with u :

$$\frac{d \ln F}{du} = \frac{d \ln f}{du} \Rightarrow \frac{d \ln F}{dc} \frac{dc}{du} = \frac{d \ln F}{dc} \frac{u}{c} = \frac{d \ln f}{du}$$

■ Const. independent of c, u, v, w :

$$\frac{1}{c} \frac{d \ln F}{dc} = \frac{1}{u} \frac{d \ln f}{du} = \frac{1}{v} \frac{d \ln f}{dv} = \frac{1}{w} \frac{d \ln f}{dw} = -\gamma$$

because it is probability...

$$f(u) = \sqrt{\frac{\gamma}{2\pi}} e^{-\frac{1}{2}\gamma u^2}$$

■ Average kinetic energy (expected value) :

$$\left\langle \frac{1}{2} m u^2 \right\rangle = \int_{-\infty}^{+\infty} \frac{1}{2} m u^2 f(u) du = \int_{-\infty}^{+\infty} \frac{1}{2} m u^2 \sqrt{\frac{\gamma}{2\pi}} e^{-\frac{1}{2}\gamma u^2} du = \frac{1}{2} m \sqrt{\frac{\gamma}{2\pi}} \int_{-\infty}^{+\infty} u^2 e^{-\frac{1}{2}\gamma u^2} du$$

$$= \frac{1}{2} m \sqrt{\frac{\gamma}{2\pi}} \sqrt{\frac{2\pi}{\gamma^3}} = \frac{m}{2\gamma}$$

using

eq. of state

$$= \frac{1}{2} kT$$

therefor

$$\gamma = \frac{m}{kT}$$

■ Maxwell distribution :

$$f(u) = \sqrt{\frac{m}{2\pi kT}} e^{-\frac{m}{2} u^2 / kT}$$

$$F(u, v, w) = \left(\frac{m}{2\pi kT} \right)^{\frac{3}{2}} e^{-\frac{m}{2} (u^2 + v^2 + w^2) / kT}$$

Example : Ideal gas (3)

- Maxwell distribution

Probability function $\propto e^{\{\text{kinetic energy}\}/kT}$

$$F(u, v, w) = \left(\frac{m}{2\pi kT} \right)^{\frac{3}{2}} e^{-\frac{m}{2}(u^2 + v^2 + w^2)/kT}$$

$$f(u) = \sqrt{\frac{m}{2\pi kT}} e^{-\frac{m}{2}u^2/kT}$$

- Only for Ideal gas ?

< Remember >

- Empirical knowledge is used.
Eq. of state for ideal gas

Relation between Temp. and average Kinetic energy

- This is the definition of Temp. !?

Let's consider more general case !

You may have learnt Thermodynamics

- Thermodynamics

(0th law : A.eq.B and B.eq.C \rightarrow A.eq.C)

1st law : Energy conservation

2nd law : Free energy decreases

3rd law : Entropy \rightarrow 0 at 0K

Everything is
derived from these
principles.

- Definitions

- Temperature

- Free energy

- Entropy

- equilibrium

- etc...

Temperature is kinetic
energy ?
(at equilibrium)

Temperature and equilibrium (general treatment)

Canonical ensemble (1)

- Assumption 1 (a model of system)

Num. of elements with energy e_i is n_i .

Total num. of elements (particle, molecule, etc...) : $N = \sum n_i$

Total energy : $E = \sum n_i e_i$

N and E
are given.

..., therefore

$n_i(e_i)$ is determined by the state. (**density of state**)

States with same $n_i(e_i)$ are indistinguishable.

- Assumption 2 (**most important !!**)

The state is observed so as to maximize
the way of shuffling of the density of state $n_i(e_i)$

Why !?

Canonical ensemble (2)

- the way of shuffling of the state
- To maximize W

$$W = \frac{N!}{n_1!n_2!n_3!\Lambda}$$

$$L \equiv \ln W - \lambda_1 \left(\sum_i n_i - N \right) - \lambda_2 \left(\sum_i n_i e_i - E \right) \rightarrow \max$$

Lagrange
multiplier

using the definition of W

$$L = \left(\ln N! - \sum_i \ln n_i! \right) - \lambda_1 \left(\sum_i n_i - N \right) - \lambda_2 \left(\sum_i n_i e_i - E \right)$$

$$\frac{\partial L}{\partial n_i} = \frac{\partial L}{\partial \lambda_i} = 0$$

using the Stirling's formula $\ln x! \approx x \ln x - x$

$$L = \left(N \ln N - N - \sum_i (n_i \ln n_i - n_i) \right) - \lambda_1 \left(\sum_i n_i - N \right) - \lambda_2 \left(\sum_i n_i e_i - E \right) \rightarrow \max$$

...

Canonical ensemble (3)

$$L = \left(N \ln N - N - \sum_i (n_i \ln n_i - n_i) \right) - \lambda_1 \left(\sum_i n_i - N \right) - \lambda_2 \left(\sum_i n_i e_i - E \right) \rightarrow \max$$

- Here, N and E are constant, therefore

$$\frac{\partial L}{\partial n_i} = -(\ln n_i + 1 - 1) - \lambda_1 - \lambda_2 e_i = 0$$



$$n_i = \exp(-\lambda_1 - \lambda_2 e_i) = \Lambda_1 \exp(-\lambda_2 e_i)$$

- Using $N = \sum n_i$

$$\frac{n_i}{N} = \frac{1}{\sum_i \exp(-\lambda_2 e_i)} \exp(-\lambda_2 e_i) \propto \exp(-\lambda_2 e_i)$$

- Comparing with Maxwell distribution

$$f(u) = \sqrt{\frac{m}{2\pi kT}} e^{-\frac{m}{2} u^2 / kT}$$

$$\frac{n_i}{N} = \frac{1}{\sum_i \exp(-e_i / kT)} \exp(-e_i / kT)$$

: Maxwell-Boltzmann distribution

e_i is not limited to be the kinetic energy.
kinetic energy, potential energy, etc...

Canonical ensemble (3)

$$L = \left(N \ln N - N - \sum_i (n_i \ln n_i - n_i) \right) - \lambda_1 \left(\sum_i n_i - N \right) - \lambda_2 \left(\sum_i n_i e_i - E \right) \rightarrow \max$$

- Here, N and E are constant, therefore

$$\frac{\partial L}{\partial n_i} = -(\ln n_i - 1) - \lambda_1 - \lambda_2 e_i = 0 \quad \Rightarrow \quad n_i = \frac{1}{\exp(-\lambda_1 - \lambda_2 e_i)}$$

- Using

$$\frac{n_i}{N} = \frac{1}{\sum_i \exp(-\lambda_1 - \lambda_2 e_i)}$$

Temperature is energy !
(used in physics)
in the most probable situation.

- Comparing with Maxwell-Boltzmann distribution

$$f(u) = \sqrt{\frac{m}{2\pi kT}} e^{-\frac{m}{2} u^2 / kT}$$

$$\frac{n_i}{N} = \frac{1}{\sum_i \exp(-e_i/kT)} \quad (-e_i/kT)$$

: Maxwell-Boltzmann distribution

e_i is not limited to be the kinetic energy.
kinetic energy, potential energy, etc...

Canonical ensemble (4)

■ Probability of the state of energy E : $\propto \exp(-E/kT)$

■ Probability of the state whose energy $> E$: $\propto \exp(-E/kT)$

■ "Assumption 2"

The state is observed so as to maximize
the way of shuffling of the density of state $n_i(e_i)$

The most probable
situation

Equilibrium !

Equilibrium is,,,
the most probable state !!

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Everything is
derived from these
principles.

- Definitions

- Temperature

- Free energy

- Entropy

- equilibrium

- etc...

Temperature is energy !
(at equilibrium)

Equilibrium is the most
probable state.

Definitions

□ Temperature

Empirical temp. °C : water } defined one or two temp.
°F : salt water }



T. can be shifted and scaled, so as proportional to energy !

□ Free energy

Helmholtz $F = U - TS$

Gibbs $G = U + PV - TS$

Free energy is same as the energy in physics ?

Why they need to be defined as above ?

□ Entropy

What is this ?

□ Equilibrium

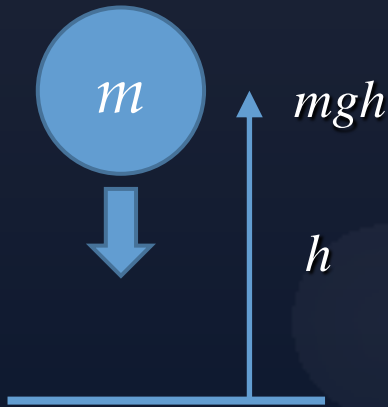
The **most probable state** ! So, we can observe always...

Free energy and entropy

Prediction of phenomena

Physicists (kinetics)

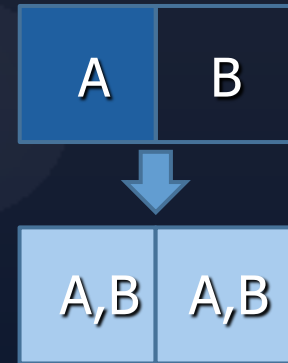
- as to decrease **Potential Energy**



Chemists (chemistry)

- as to decrease **Free Energy**

Ideal gases



After the phenomena

Decreased
Not changed
? (Not changed)

Potential Energy
Total energy
Free Energy

Not changed
Not changed
Decreased

Chemists wanted to introduce **something like potential energy**, to predict the phenomena.

Chemists wanted to introduce **something like potential energy**, to predict the phenomena.

The **something**; related to the work or energy, used in thermodynamic technology, (ex. steam engine, chemical reaction, etc...)

Definition of free energy :

Energy that can be converted into a work in a process of

constant temperature and volume (Helmholtz). $F = U - TS$

constant temperature and pressure (Gibbs). $G = U + PV - TS$

(Two of three T , P , & V determine its state.)

Energy or work in kinetics

What's this ! ??

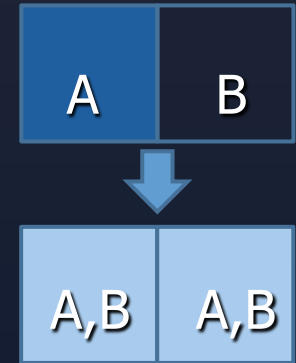
$$G = U + PV - TS$$

Definition of free energy :

Energy that can be converted into a work in a process

$$G = U + PV - TS$$

What's this ! ??



- The TS term increases in the mixing process, even if kinetic and potential energy never changes.
- The mixing process never goes back.

Increase in TS term \rightarrow Decrease of G

= Increase of the usable energy

= Degrade of energy

Definition of free energy :

Energy that can be converted into a work in a process

$$G = U + PV - TS$$

What's this ! ??

- There exists a **part of energy**, which we can not take as a work.
- **It** is proportional to the temperature.

The **degrade** is proportional
to the temp...

Degrade of the
energy exits...

Empirical knowledge
of chemists

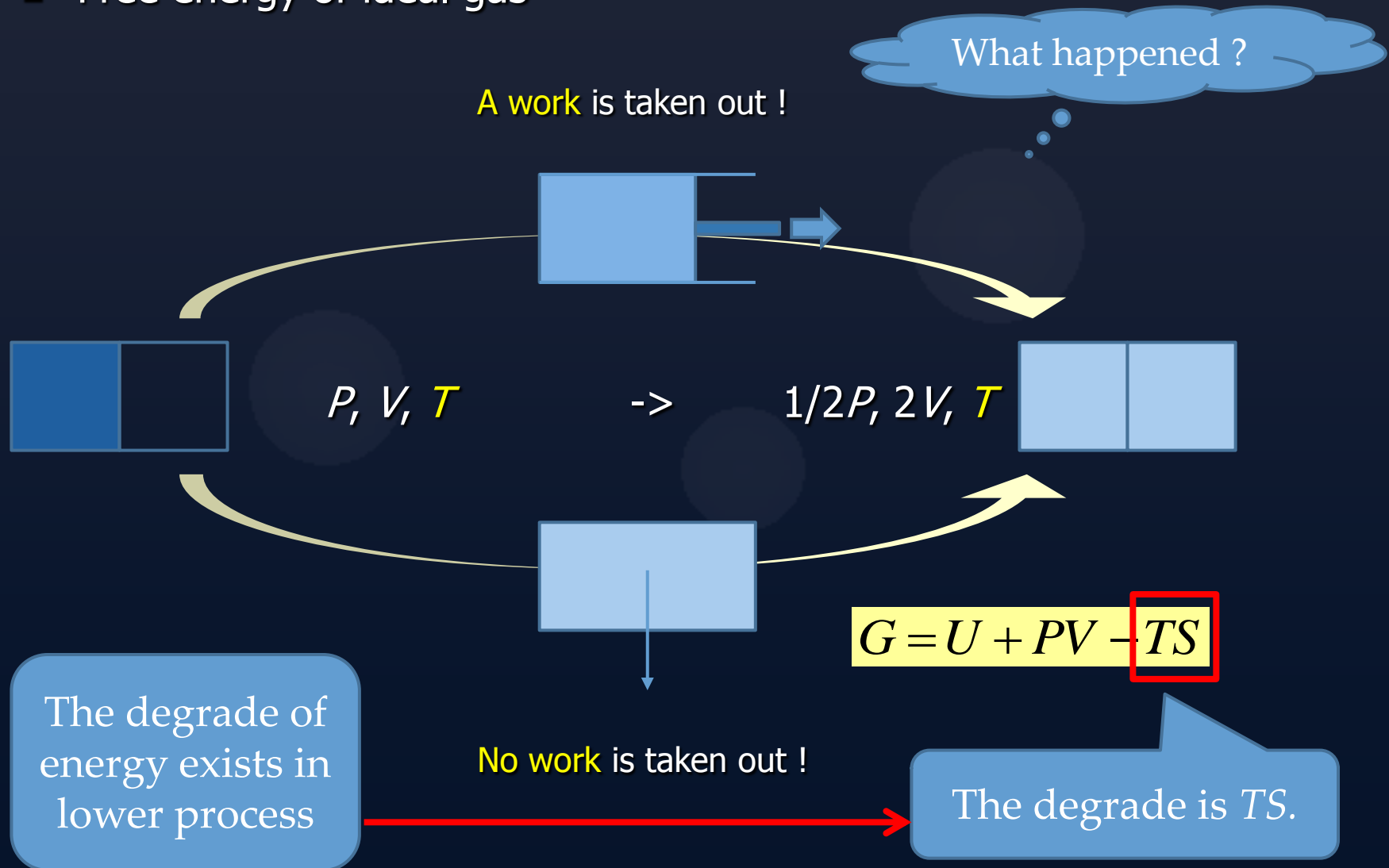
- S is defined
as the **entropy**.

What's the entropy ! ??

This is the definition of entropy **in thermodynamics**.

Boltzmann's entropy formula & Free energy

■ Free energy of ideal gas



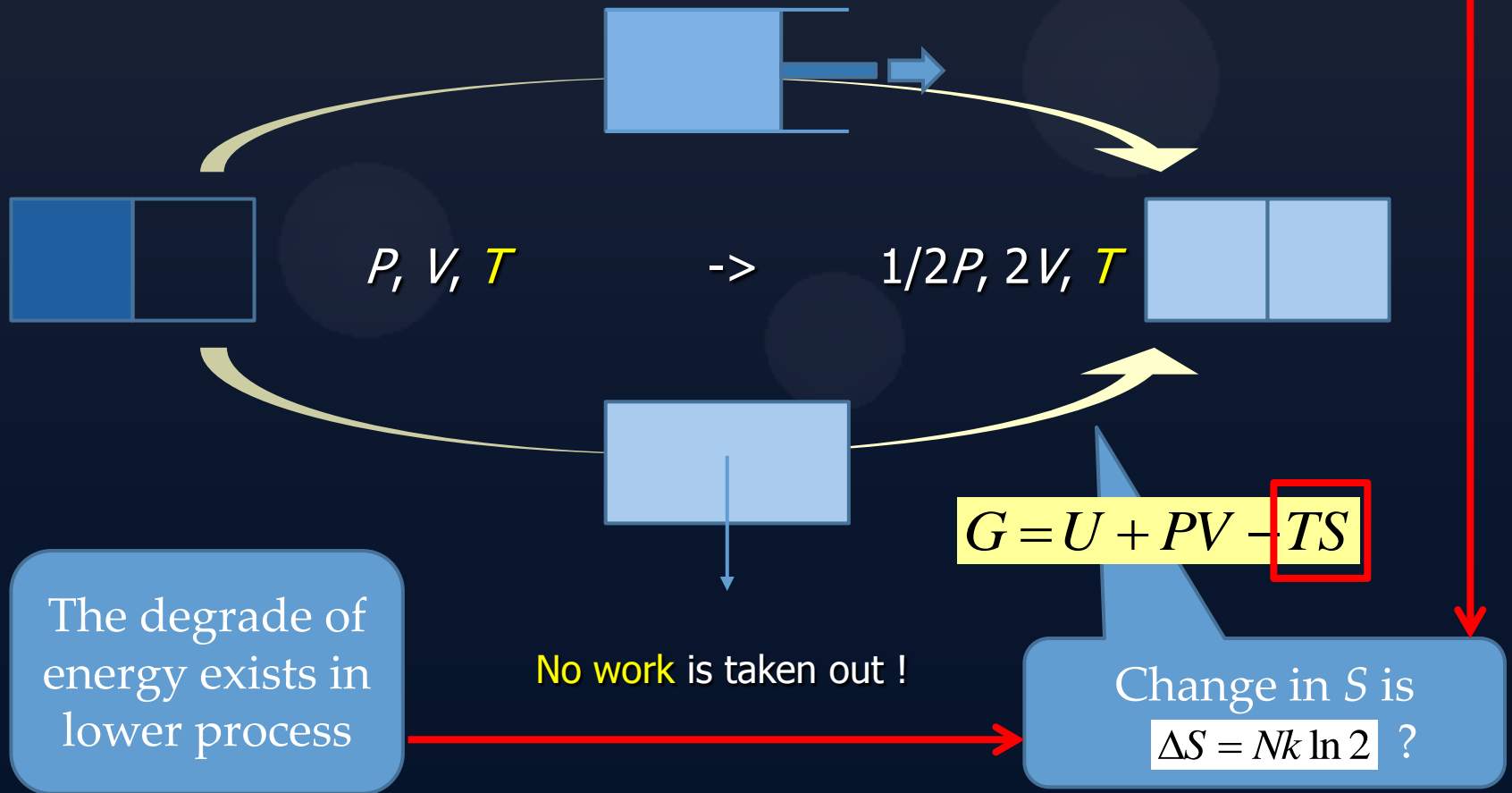
Boltzmann's entropy formula & Free energy

■ Free energy of ideal gas

$$PV = NkT$$

A work is taken out !

$$\int_V^{2V} PdV = \int_V^{2V} \frac{NkT}{V} dV = NkT \ln 2$$



Statistical Mechanics and Thermodynamics

- Thermodynamics
(0th low : A.eq.B and B.eq.C -> A.eq.C)
1st low : Energy conservation
2nd low : Free energy decreases
3rd low : Entropy -> 0 at 0K

Boltzmann's entropy formula : $S = k \ln W$



Grave-stone of Boltzmann...

- Statistical Mechanics
Ergodic theorem : time average = space average, (in equilibrium)

Boltzmann's entropy formula & Free energy

■ W : Number of state $S = k \ln W$

v : small control volume

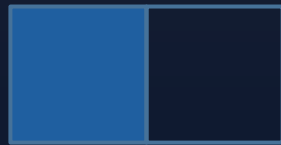
V/v : num. of the c. volume

$2V/v$: num. of the c. volume

Num. of the way to shuffle to distribute the atoms to the each c. volume

$$W_0 = (V / v)^N$$

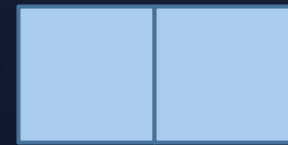
$$W = (2V / v)^N = W_0 2^N$$



P, V, T

\rightarrow

$1/2P, 2V, T$



$$S_0 = k \ln W_0$$

$$\begin{aligned} S &= S_0 + \Delta S \\ &= k \ln(W_0 2^N) \\ &= S_0 + Nk \ln 2 \end{aligned}$$

No work is taken out !

$$\Delta S = Nk \ln 2$$

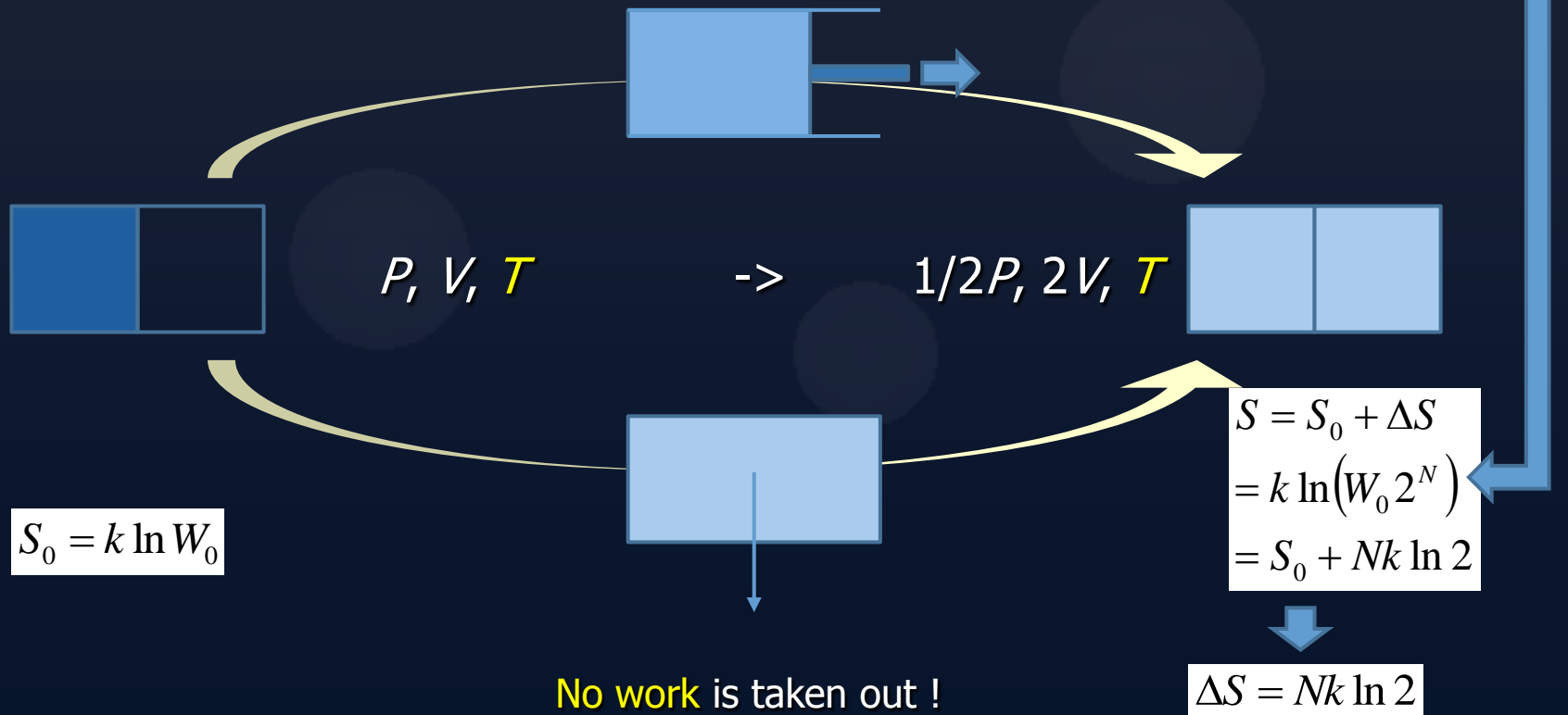
Boltzmann's entropy formula & Free energy

Free energy of ideal gas

$$PV = NkT$$

A work is taken out !

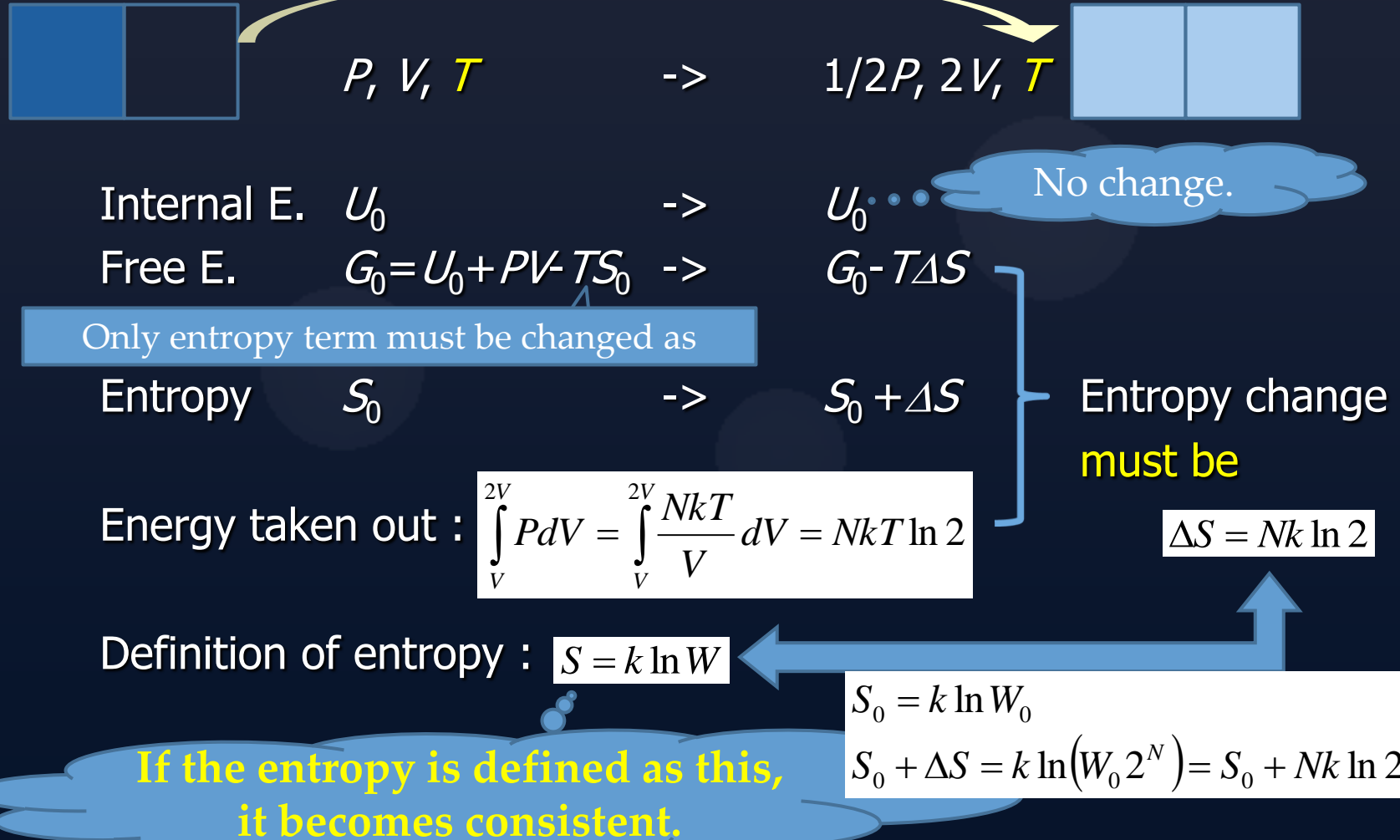
$$\int_V^{2V} PdV = \int_V^{2V} \frac{NkT}{V} dV = NkT \ln 2$$



$$\Delta S = Nk \ln 2$$

Boltzmann's entropy formula & Free energy

Free energy of ideal gas



Boltzmann's entropy formula & Free

Check of
understanding

■ Free energy of ideal gas

$$PV = NkT$$

A **work** is taken out !

$$\int_V^{2V} P dV = \int_V^{2V} \frac{NkT}{V} dV = NkT \ln 2$$

We can take out the infinite work ! ???...
although internal energy must be finite.

P, V, T

->

$1/2P, 2V, T$

W depends on volume ! ???...
Energy **never** depend on volume...

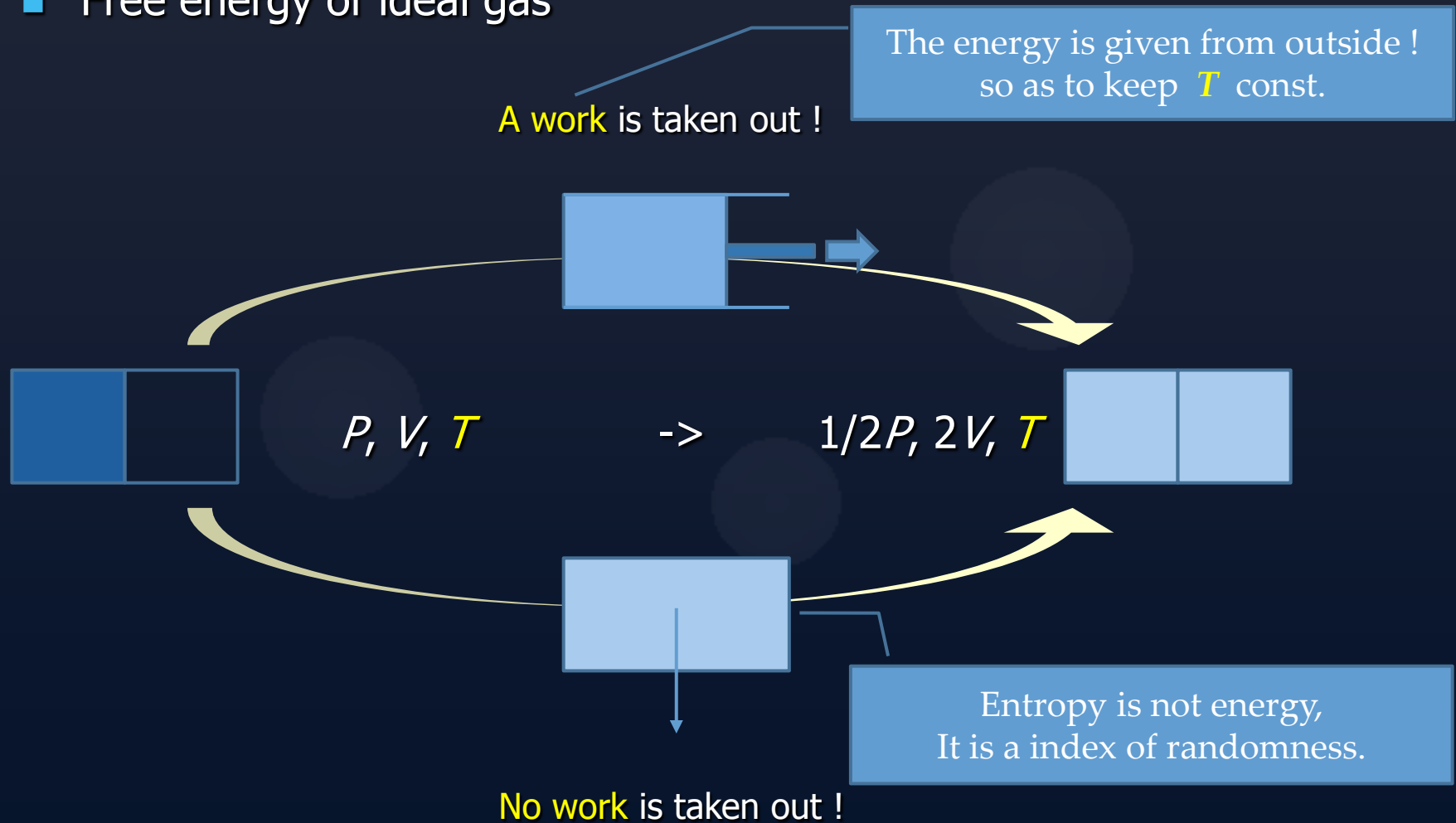
No **work** is taken out

$$S_0 = k \ln W_0$$

$$S_0 + \Delta S = k \ln(W_0 2^N) = S_0 + Nk \ln 2$$

Boltzmann's entropy formula & Free energy

■ Free energy of ideal gas



Introduction of Free energy & entropy & temperature

- Definition of **Free Energy** :

Energy that can be converted into a work.

(There exist the energy part we can not take out.)

Internal energy
= Kinetic energy

we can take out from outside
(by heat transfer...)

$$G = U + PV - TS$$

Degrade of the
energy...

- Definition of **Entropy**

Related to the randomness

$$S = k \ln W$$



Definition of **Temperature**

□ Temperature

Empirical temp. °C : water } defined one or two temp.
°F : salt water



T. can be shifted and scaled, so as proportional to energy !

□ Free energy

Helmholtz $F = U - TS$

Gibbs $G = U + PV - TS$



$$G = U + PV - TS$$

Free energy is energy that can be converted into work.
If defined as above, it become consistent.

□ Entropy

$S = k \ln W$ corresponding to number of states, i.e. randomness.

□ Equilibrium

Most probable state ! So, we can observe always...

These are the interface between
Thermodynamics and statistical mechanics

Statistical Mechanics and Thermodynamics

- Thermodynamics
(0th low : A.eq.B and B.eq.C -> A.eq.C)
1st low : Energy conservation + ...
2nd low : Free energy decreases
3rd low : Entropy -> 0 at 0K

Boltzmann's entropy formula : $S = k \ln W$

- Statistical Mechanics
Ergodic theorem : time average = space average, (in equilibrium)



Grave-stone of Boltzmann...

Based of statistical mechanics,
Thermodynamics can be utilized !

Activation energy and Arrhenius plots and rate-determining step

- chemical reaction
- diffusion coefficient
- others... (evaporation, adsorption, etc...)

Arrhenius plots is to know the mechanism of rate-determining step.

Specific heat capacity (1: ideal gasses)

- Remember thermodynamics ! (How have you learnt ?)

$$C_v = \frac{1}{N} \left(\frac{\partial U}{\partial T} \right)_v \quad \text{why ?} \quad C_p = \frac{1}{N} \left(\frac{\partial H}{\partial T} \right)_p \quad \text{why ?}$$

- Definition

Heat required to increase unit temp.
- under constant volume

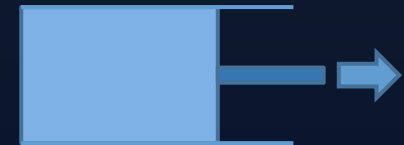
$$\Delta Q = \Delta G = \Delta U = NC_v \Delta T$$

- under constant pressure

$$\Delta Q = \Delta G = \Delta U + P\Delta V - \Delta T\Delta S = \Delta U + P\Delta V = NC_p \Delta T$$

- Mayer's relation

$$C_p - C_v = k$$



$$NC_v \Delta T + P\Delta V = NC_p \Delta T$$

$$G = U + PV - TS$$

Specific heat capacity (2: metals)

- Einstein's model

- **Dulong-Petit's Low**

- Debye model

What is the U ?

- Ionic cores



$$U = \langle e_k \rangle + \langle e_p \rangle = 3\frac{1}{2}kT + 3\frac{1}{2}kT = 3kT$$

- Electrons = negligible



Quantum statistical mechanics

$$\frac{\Delta Q}{\Delta T} = 3k$$

Phase diagram

- Stability of phases
- Phase rule
- Phase diagram
- Calculation of Phase Diagram (CALPHAD)